

# How (and Why) Stars Die

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Twinkle, twinkle, little star  
How I wonder what you are ...



# Stars Are Distant Suns!



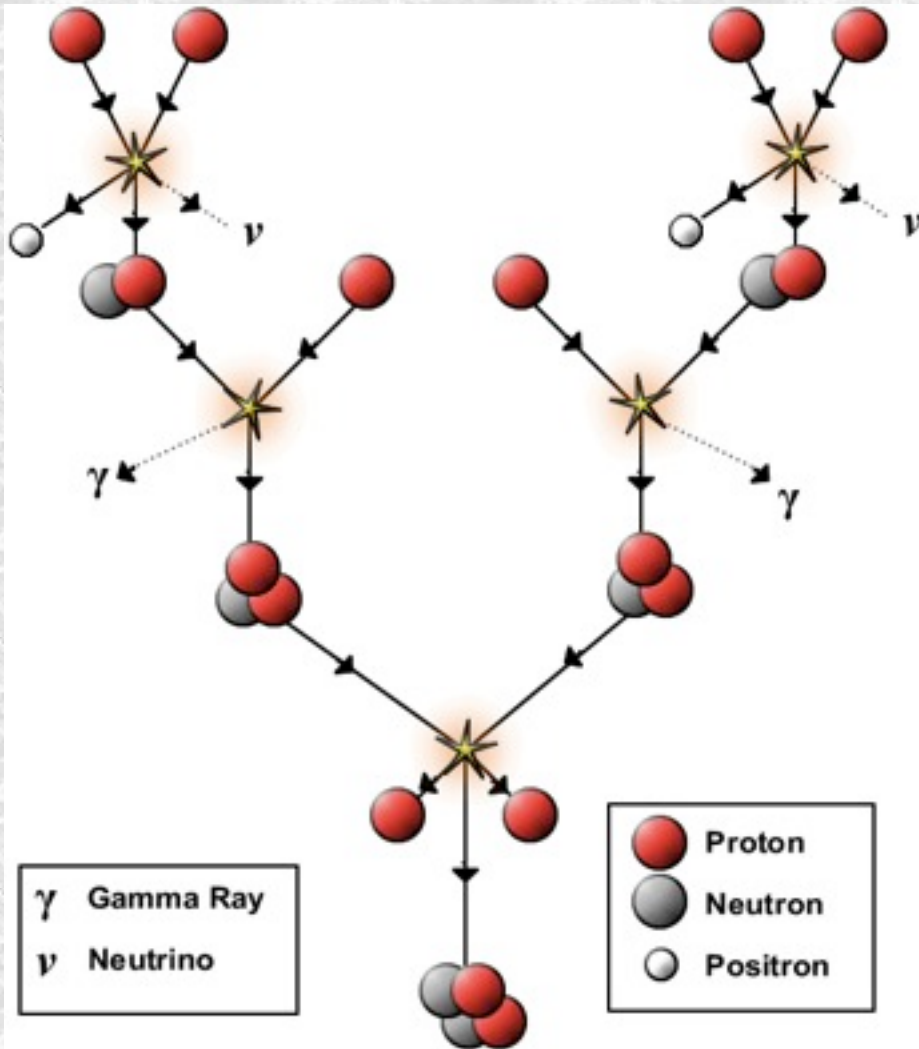
- The sun is a million times bigger than the earth
- Stars shine by producing energy; moons and planets shine by reflected light
- Stars produce energy in their hot core, where the temperature is 15,000,000 degrees!

# The stars are born in **nebulas**: clouds of gas and dust in space



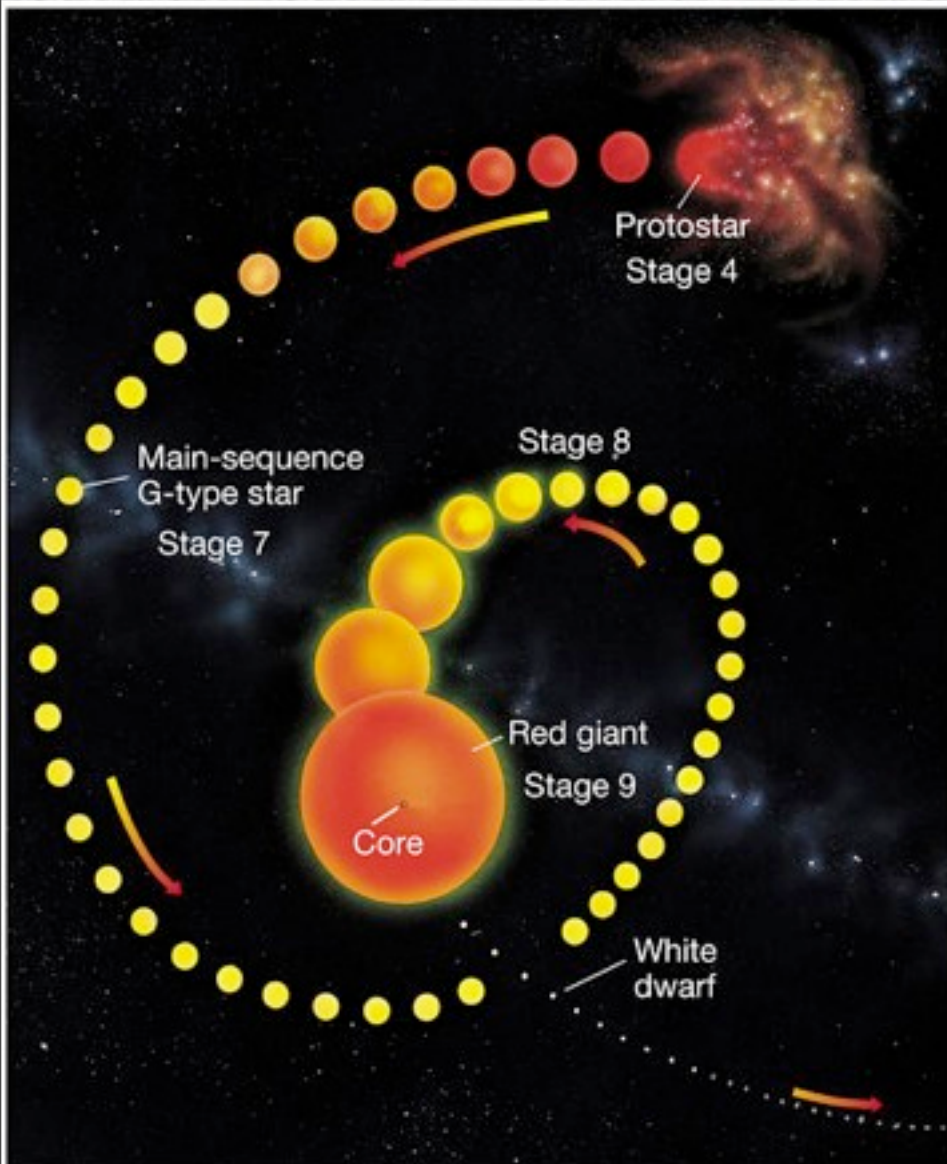
- Nebulas, like the Orion Nebula which you can see with binoculars, are clouds of gas and dust in space
- The gas and dust comes from the birth of the universe, and also from dying stars

# The stars make energy by **nuclear fusion** of hydrogen



- The stars are made of  $\frac{3}{4}$  hydrogen and  $\frac{1}{4}$  helium
- In their hot, dense cores, the hydrogen fuses into helium: “thermonuclear fusion”
- If we could harness this process on earth, it would solve many of our energy problems!

# Stars have long lifetimes!



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- The sun's power is 400 trillion, trillion Watts
- Still, it has enough fuel to last for 10 billion years
- Most stars have even longer lifetimes – trillions of years
- Only rare, more massive stars have shorter lifetimes – millions of years

# But the stars eventually run out of fuel



- Every energy supply runs out eventually
- The stars run out of energy when all of the hydrogen in their hot core is changed into helium
- The core shrinks, to try and squeeze some energy out of the helium
- The rest of the star swells into a **red giant**

# The red giant begins to pulse

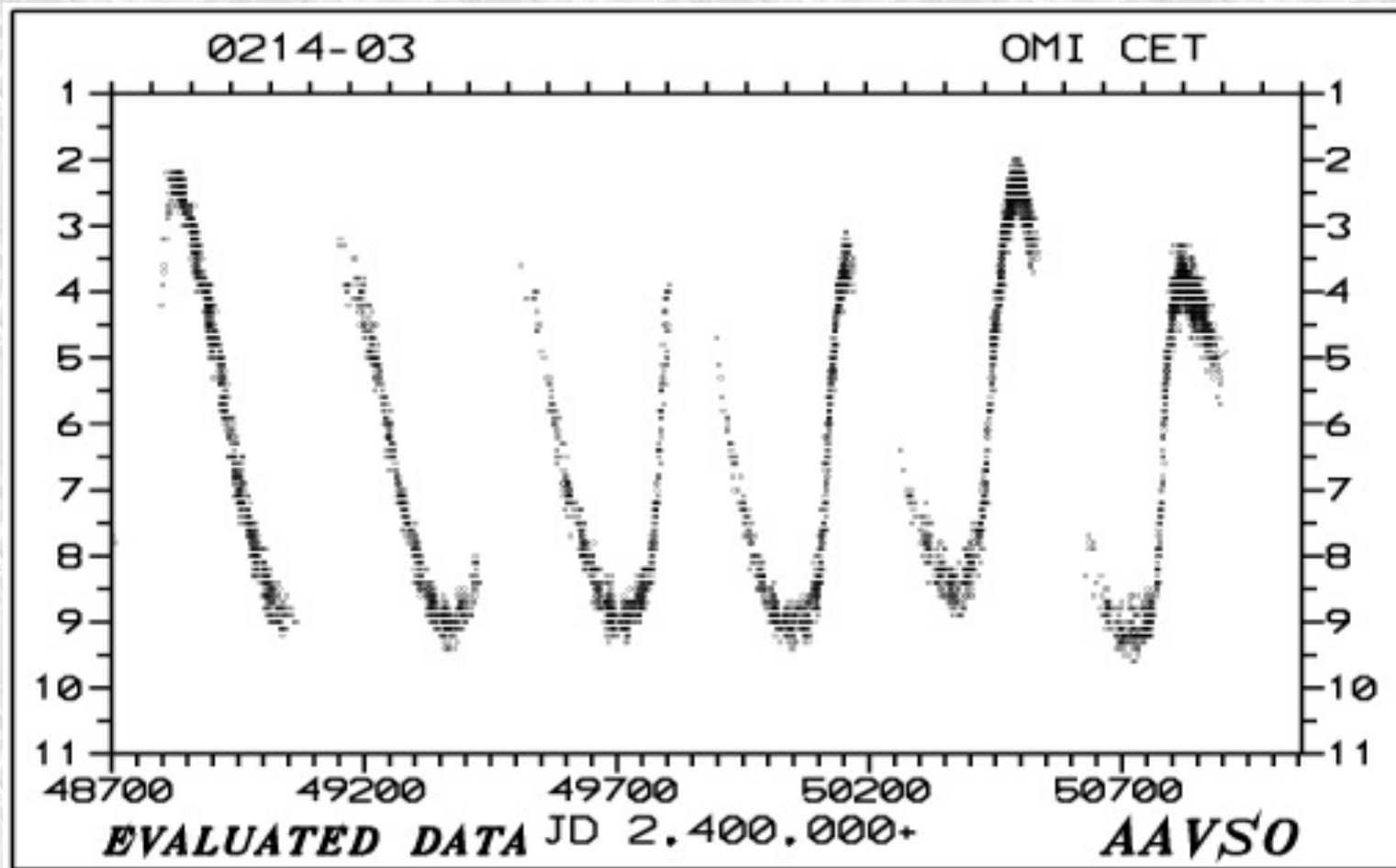


- The red giant swells, cools and brightens, swallowing the inner planets
- The outer layers become unstable
- They begin to pulse, in and out, every few weeks or months
- The pulsing drives off the outer layers into space, forming a **planetary nebula**

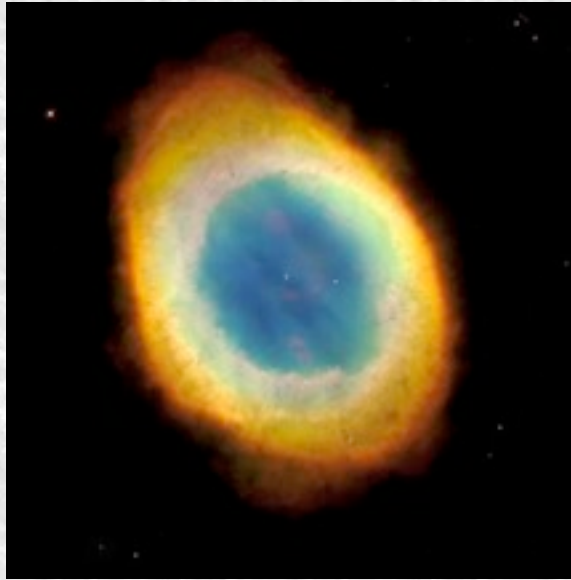


# Mira Ceti is a **Pulsing Red Giant**

Every 330 days, it brightens and fades as it expands and contracts. These measurements were made by skilled amateur astronomers



# Dying Suns – Planetary Nebulas

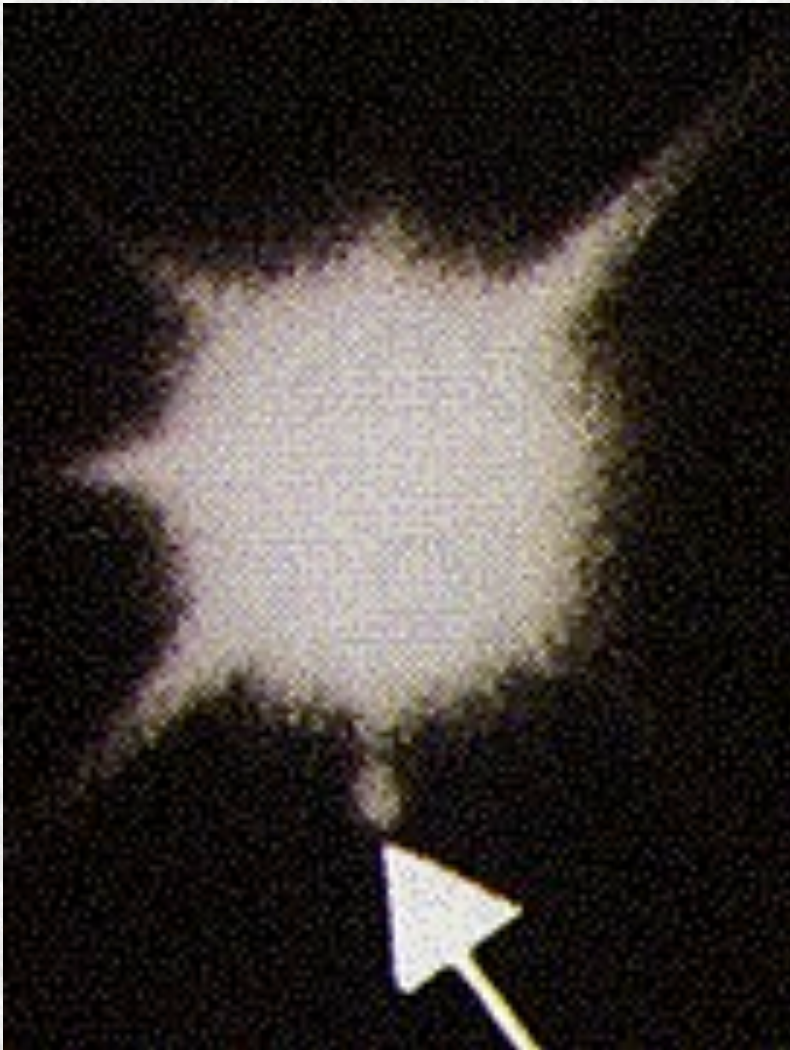


# White Dwarfs



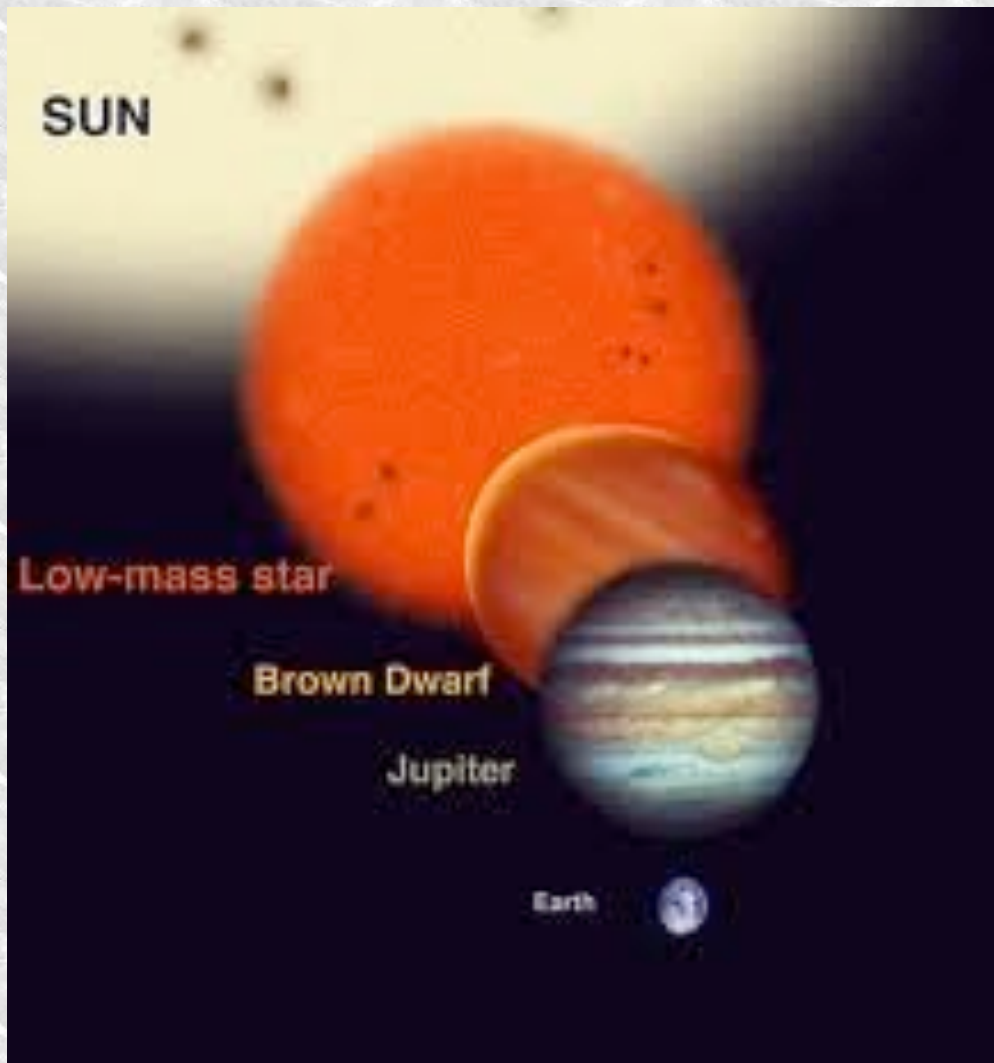
- A **white dwarf** is the shrunken corpse of a normal star like the sun
- It has the mass of a star, in the volume of the earth
- It's density is a million times that of water
- It has no energy; it just cools off like a cinder in a fire

# Sirius B: A Nearby White Dwarf



- Sirius is the brightest star in the night sky
- It is actually a pair of stars, orbiting each other
- One (Sirius A) is a normal star; the other (Sirius B) is a white dwarf
- It is the remains of a star that has already lived and died

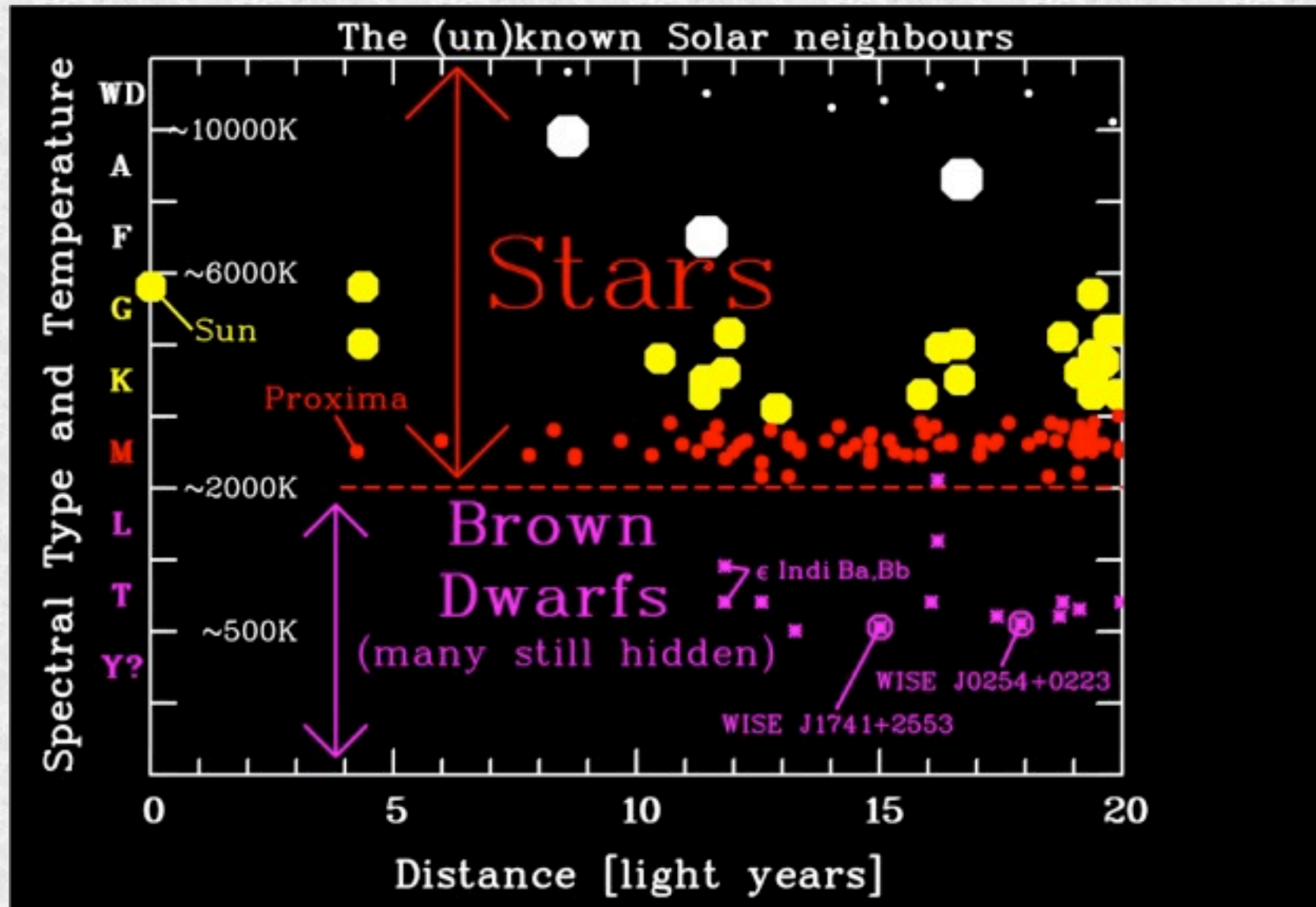
# The sun is not an average star!



- You may have learned that the sun is an average star
- It's not; it's bigger and more powerful than 90% of other stars
- Most stars are like the “low mass” star at left
- Massive stars are very rare!

# The sun is not an “average” star!

Smaller stars are common; larger stars are rare



Institute of Astrophysics, Potsdam

# When massive stars run out of fuel



- Rare, massive stars fuse hydrogen into helium, carbon, oxygen, and elements as heavy as iron
- When they run out of fuel, their cores collapse violently under their own weight
- The gravitational energy release explodes the star: a **supernova**

# Brightest Supernova in 400 Years!

This supernova was discovered by University of Toronto astronomer Ian Shelton on February 23-24, 1987 – 25 years ago!





# Supernova remnants



- The supernova remnant is blasted into space at thousands of km/sec!
- This material has been enriched in the elements which were created in the star by nuclear fusion
- The material forms new nebulas from which new stars and planets and life are made
- **You are starstuff!**

# Neutron Stars!



- The core of a rare, massive star collapses under its own weight until it is a ball of neutrons – a **neutron star**
- Its density is a million tonnes per  $\text{cm}^3$
- It can spin faster than a kitchen blender!
- It emits pulses of radiation as it spins; it is a **pulsar**

# When a *very* large star runs out of fuel

- One star in billions is so massive that, when it runs out of fuel, and collapses under its own weight, its core becomes a **black hole**
- Its density is so great that its gravity is so strong that nothing can escape from it – not even light

# How the first black hole was discovered

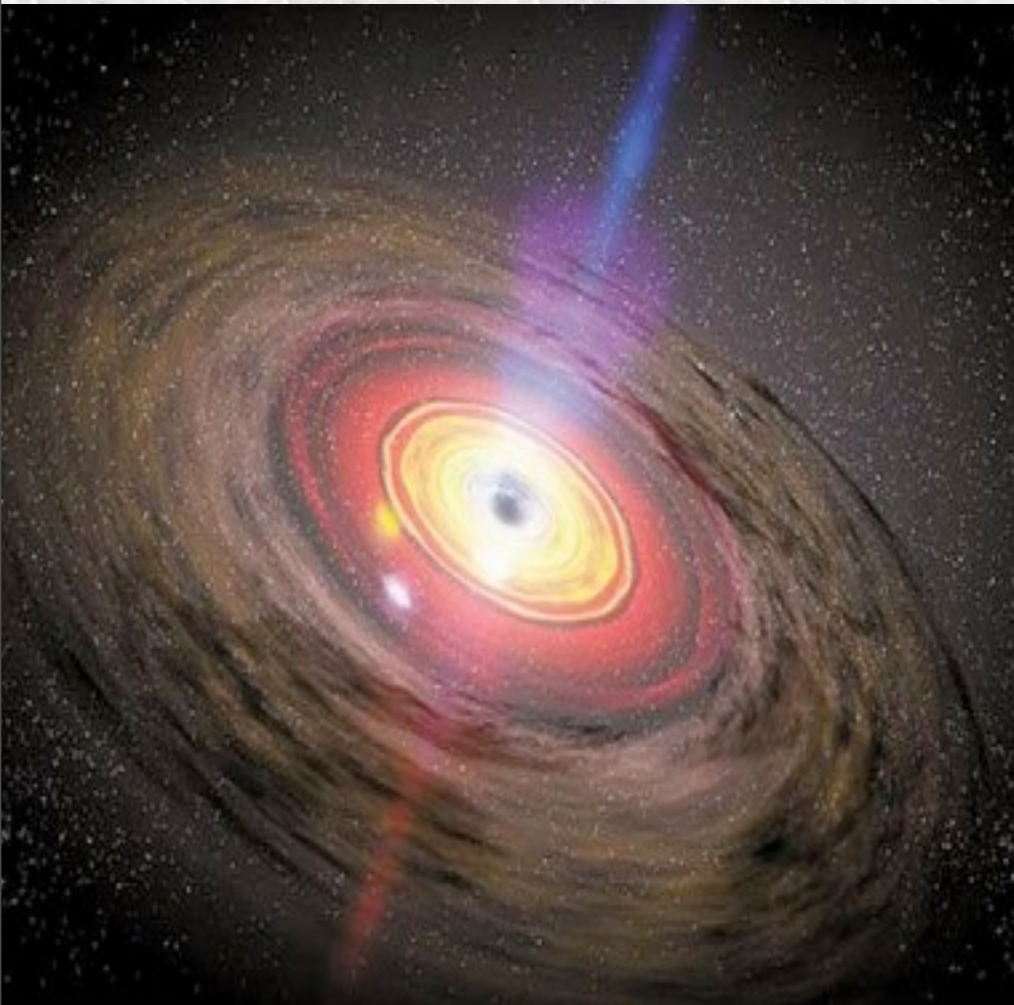


- By University of Toronto astronomer Tom Bolton at the Dunlap Observatory in Richmond Hill ON
- By the black hole's pull on a normal star going around it
- And from X-rays produced as gas falls into the black hole

# Myths about black holes

- You can see a black hole. **No, light can't escape.**
- The gravity of a black hole is different from normal gravity. **No, same gravity.**
- The sun – and all other stars – will turn into a black hole. **No, just very rare, massive stars.**
- Black holes are giant cosmic vacuum cleaners that swallow everything around them. **No, only things very nearby, such as gas from a star which is orbiting them.**
- Matter (such as you) that falls into a black hole will appear somewhere else in the universe. **No, it stays right there.**

# Supermassive black holes



- Supermassive black holes, millions of times more massive than the sun, are found at the centre of galaxies such as our Milky Way
- We don't know how they formed, so we are looking in young galaxies that are just forming

# How astronomers know all this



- We observe and study the stars with telescopes on the ground and in space
- We use scientific knowledge to understand how the stars work

# Find out more

- The University of Toronto astronomy outreach website:
- [universe.utoronto.ca](http://universe.utoronto.ca)
- My Education and Outreach webpage:
- [www.astro.utoronto.ca/~percy/EPOindex.htm](http://www.astro.utoronto.ca/~percy/EPOindex.htm)
- Canadian Astronomical Society:
- [www.cascaeducation.ca](http://www.cascaeducation.ca)